## Determining the ratio $\frac{r_p}{r_e}$ for the outer planets.

In Figure 1 we consider the triangular configuration of the sun, earth, and a planet at a particular moment in time,  $t_1$  in a heliocentric model. The distance to the earth from the sun is  $r_e$  and the distance from the sun to the planet is  $r_p$ . Suppose at some time in the year the earth the planet and the sun are in conjunction. Imagine this line SEP is extended to the Celestial Sphere, indicated by a star  $S_1$ . At conjunction define  $t_0 = 0$ , and at some later time,  $t_1$  the earth has moved to a new position making an angle with respect to the direction to  $S_1$  by an angle  $\alpha$ 

$$\alpha = 360^{\circ} \frac{t_1 - t_0}{T_e} \tag{1}$$

At the same time the planet has moved through an angle  $\beta$ 

$$\beta = 360^{\circ} \frac{t_1 - t_0}{T_p} \tag{2}$$

Here  $T_e$  and  $T_p$  are the orbital periods of the earth and the planet. At the vertex of the sun we determine the angle  $\gamma$ 

$$\gamma = \alpha - \beta \tag{3}$$

Now at time  $t_1$  we make an observation of the angle,  $\theta$ , of the planet with respect to the direction of  $S_1$ . We can now determine the angle  $\phi$  at the vertex of the earth

$$\phi = 180 - \alpha - \theta \tag{4}$$

Hence, we know from our studies of Euclidian geometery that angle-side-angle information is enough to determine the other parameters of the triangle  $\Delta SEP$ . In particular this will give us the ratio  $\frac{r_p}{r_e}$ . A plane figure drawing with a protractor and ruler would be one way to do this. Another way is to do a calculation, for example. In figure 1 we draw in a perpendicular line h from the vertex e to the  $r_p$  line. The angle  $\delta$  at the vertex p is

$$\delta = 180 - \gamma - \phi. \tag{5}$$

$$h = stan(\delta) \tag{6a}$$

$$h = (r_p - s)tan(\gamma) \tag{6b}$$

$$s = \frac{r_p tan(\gamma)}{tan(\delta) + tan(\gamma)} \tag{6c}$$

We combine equations 6a and 6b to obtain a statement for s in equation 6c. From the vertex at s we recognize that  $cos(\gamma) = \frac{rp-s}{r_e}$  We can eliminate s from this equation with equation 6c. The result is

$$\frac{rp}{re} = \left(\frac{\tan(\delta) + \tan(\gamma)}{\tan(\delta)}\right)\cos(\gamma) \tag{7}$$



Figure 1: Triangular configuration of the sun and earth and a planet,  $\Delta SEP$  at time  $t_1.$